

What can we learn from the energy dependence of strangeness production at RHIC?

J. Speltz^a for the STAR Collaboration

Institut Pluridisciplinaire Hubert Curien, 23 rue du Loess, BP 28, Strasbourg, France

Received: 3 August 2006 /

Published online: 30 November 2006 – © Springer-Verlag / Società Italiana di Fisica 2006

Abstract. Measurements of strange and multi-strange particles with the STAR detector at center of mass energies per nucleon pair ($\sqrt{s_{NN}}$) of 62.4 GeV and 200 GeV in ultra-relativistic Au+Au collisions at RHIC are presented. The results are discussed in order to get insight into chemical and dynamical properties of the created medium. The former are obtained by comparison of transverse momentum (p_T) integrated observables such as yields and particle ratios to statistical models, while the latter use p_T dependent quantities such as single particle spectra and elliptic flow measurements to challenge hydrodynamical model calculations and parameterizations. The discussion is orientated towards the energy dependence of these properties by confronting the results at the different RHIC energies but also with the lower SPS energies.

1 Introduction

Strangeness observables have been suggested, since the beginning of the study of relativistic heavy ion collisions, to yield possible effects of the creation of a quark gluon plasma [1]. These observables are supposed to give valuable insight into the properties of the created medium, as all strange valence quarks in the final state originate from the collision and are not present in the incoming nuclei. The study of these features in terms of chemical and kinetic freeze-out conditions by comparing the data to predictions of models may then reveal different aspects and characteristics of the fireball they arise from. The examination of the energy dependence of these properties, especially with the results at the intermediate energy of 62.4 GeV that stands between the top RHIC and the SPS energies, gives the possibility to access a more global picture of the characteristics of the medium, such as a possible onset of the formation of a QGP.

The presented STAR [2] data are from Au + Au collisions delivered by RHIC and were measured with its main tracking device, the time projection chamber (TPC) [3]. The 200 GeV results are from the RUN II [4] and the 62.4 GeV results come from the RUN IV [5, 6]. The results are for mid-rapidity ($|y| < 1$) if not otherwise stated.

2 Chemical properties

The relative abundance of the different particle species and consequently the chemical composition of the medium

is frozen at chemical freeze-out which marks the end of inelastic interactions among the particles. The temperature, T_{ch} , that characterizes this freeze-out can be accessed by statistical models [7–11]. These models compute the production rates of the different particles by using a limited set of parameters and assume a statistically equilibrated matter. Hence one can consider that these parameters allow the characterization of the chemical properties of the system. In addition to T_{ch} , these parameters are the chemical potentials of light quarks μ_q ($q = u, d$; baryon chemical potential $\mu_B = 3\mu_q$) and of strange quarks μ_s , as well as the strange quarks phase-space occupancy factor, γ_S , that characterizes the strangeness saturation. For a given collision energy and system, these parameters are obtained by adjusting measured particle ratios.

In Fig. 1 (top panel) particle ratios obtained from 62.4 GeV central Au + Au data (circles) are shown with results from statistical model calculations obtained by adjusting data ratios involving π^\pm , K^\pm , p and \bar{p} , Λ and Ξ (lines) with a $\chi^2/ndf = 2.3/5$. The results on the Ω/π and $\bar{\Omega}/\Omega$ ratios are then a pure prediction from the statistical model. The difference between data and model prediction is evaluated by the number of standard deviations as shown in the bottom panel of Fig. 1. A comparable figure, exposing the good agreement between data and statistical model calculations at 200 GeV, can be found elsewhere [12].

Figure 2 shows the dependence of T_{ch} and γ_S on the mean number of participant nucleons, $\langle N_{part} \rangle$, determined by a Glauber model calculation [13]. The closed circles are for 62.4 GeV and the open circles are for 200 GeV. The results at both energies were obtained with the same model [11]. The temperature is constant over all centralities at ~ 160 MeV, close to the critical temperature of $T_c \sim 170$ MeV obtained from lattice QCD calculations [14].

^a e-mail: jeff.speltz@ires.in2p3.fr

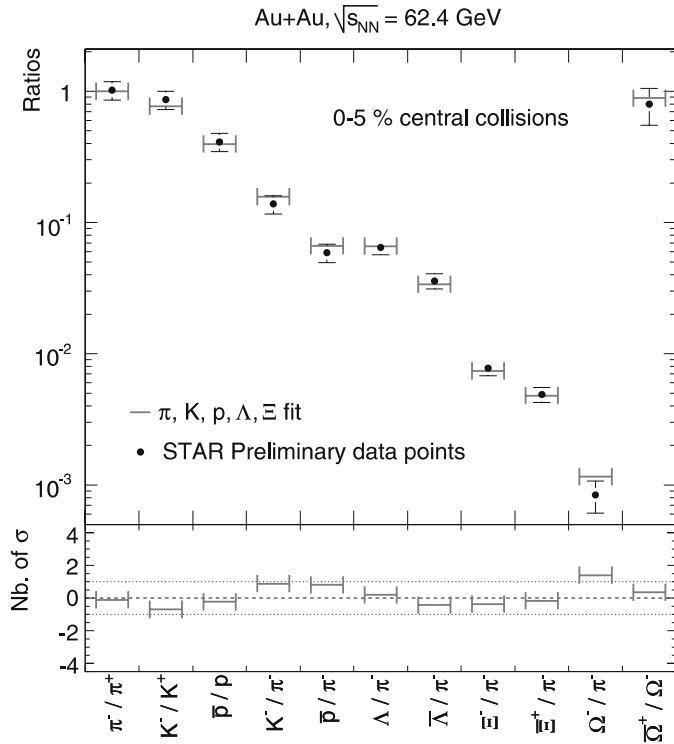


Fig. 1. *Top panel:* “Stable” particle ratios obtained by measurements for most central collisions at $\sqrt{s_{NN}} = 62.4$ GeV (*points*) superposed with statistical model calculations using [11] (*lines*). The experimental data have statistical and systematic errors added quadratically. *Bottom panel:* Number of standard deviations between data and model calculations

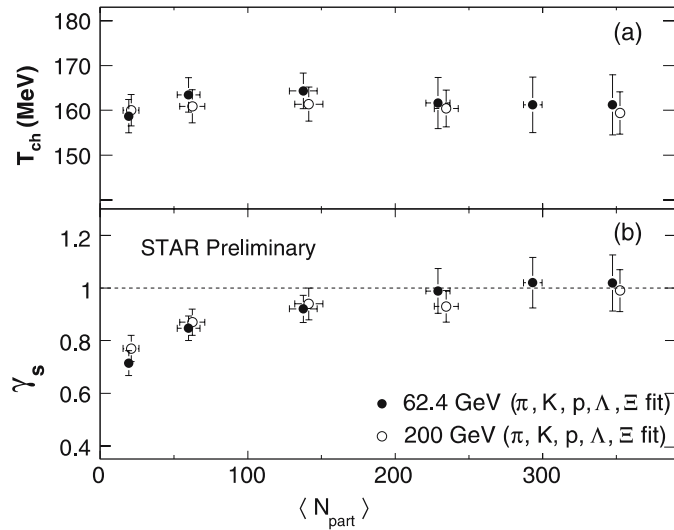


Fig. 2. Centrality dependence ($\langle N_{part} \rangle$) of **a** chemical freeze-out temperature T_{ch} and **b** strangeness saturation factor γ_s for 62.4 GeV (*closed circles*) and 200 GeV (*open circles*). Statistical and systematic errors are added quadratically

In contrast, γ_s monotonously rises from a around 0.7 for peripheral collisions to saturate at a value close to unity for the most central collisions. This indicates the presence of a state of matter where strangeness production is

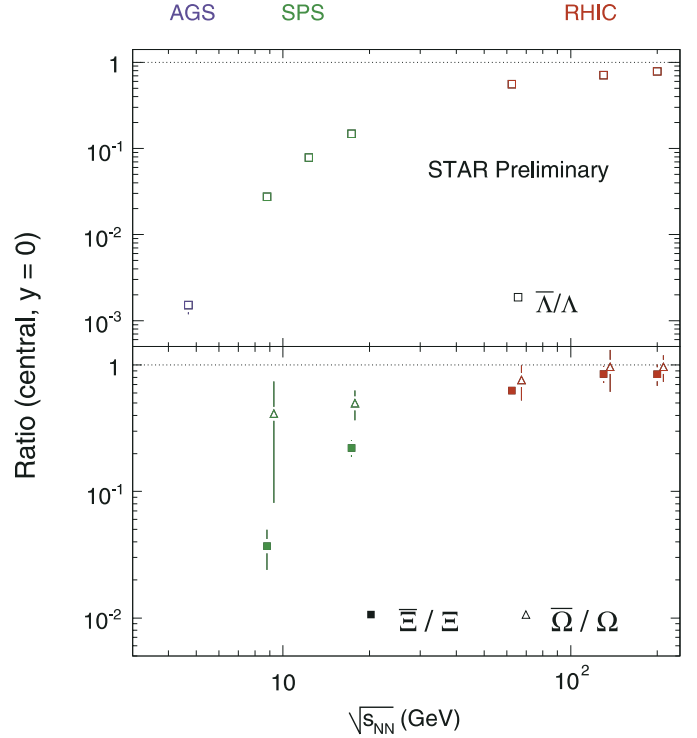


Fig. 3. Energy dependence ($\sqrt{s_{NN}}$) of $\bar{\Lambda}/\Lambda$, $\bar{\Xi}/\Xi$ and $\bar{\Omega}/\Omega$ ratios for central nucleus-nucleus collisions at mid-rapidity [18–22]

close to saturation and may thus reveal the achievement of a global chemical equilibrium for strangeness. The results for 200 GeV shown on the same figure as open circles display the same behaviour and magnitude within error bars as the 62.4 GeV data, while the SPS data at the lower energy of 17.3 GeV may give a slightly lower value for γ_s [9], with a comparable centrality dependence.

The baryon chemical potential obtained at 62.4 GeV ($\mu_B = 87 \pm 13$ MeV) is higher as at 200 GeV ($\mu_B = 24 \pm 4$ MeV) but stands where expected to be by interpolation between SPS and RHIC results [15]. This variation in μ_B in statistical models results directly from the the anti-baryon to baryon ratios in data. In Fig. 3 the evolution of this ratio for Λ , Ξ and Ω is represented as a function of $\sqrt{s_{NN}}$. The smooth rise in these ratios and the approach to unity for top RHIC energy translates the transition from a transport dominated regime to a dominance of quark anti-quark pair production. The latter leads to a baryon-free environment equivalent to lower baryon chemical potential.

The astonishing success of statistical models over a large range of energies, especially also for $p+p$ and $e^- + e^+$ collisions [16, 17] requires a word of caution in the interpretation with respect to a chemical equilibration of the system. It is possible that it only looks as if the system follows statistical laws because the experimental ratios result from averaging over a large number of events. The abundances from a single event may then not necessarily be reproducible by models assuming chemical equilibrium.

3 Dynamical properties

Dynamical properties of the system created in relativistic heavy-ion collisions can be accessed by p_T dependent observables, such as particle spectra and differential elliptic flow [23]. The comparison of data with models may help getting information on these properties, such as the temperature T_{dec} (T_{kin}) of the kinetic freeze-out when elastic interactions end and the magnitude of collective motion during the expansion. In turn, these should give insight into the validation or not of the hypothesis of local thermalization and even the creation of a de-confined phase of quarks and gluons which should be described by a dedicated equation of state.

At first, we extract dynamical parameters with the hydrodynamically inspired, so-called *Blast-Wave* parameterization [24, 25], by performing a direct fit on the spectra. The two free parameters in the fit, T_{kin} and $\langle\beta_T\rangle$ are interpreted as the kinetic freeze-out temperature and the mean transverse flow velocity, respectively. Figure 4 shows the confidence-level contours in $T_{\text{kin}} - \langle\beta_T\rangle$ space obtained by *Blast-Wave* fits on the spectra of different particle species at different energies. The 62.4 GeV and 200 GeV contours reveal a clear discrepancy between the transverse flow of the light particles (π^\pm , K^\pm , p and \bar{p}) and the multi-strange baryons (Ξ and Ω). For the light particles, a decrease in radial flow is observed for the lower energy. Such a decrease for multi-strange baryons is not easy to identify between the 200 GeV and the 62.4 GeV data points due to the uncertainties. But it may be observed when comparing to the NA57 results [26] at 17.3 GeV. A caveat is that the NA57 contour is for a much larger centrality range which includes peripheral collisions that may lead to less radial flow. A dis-

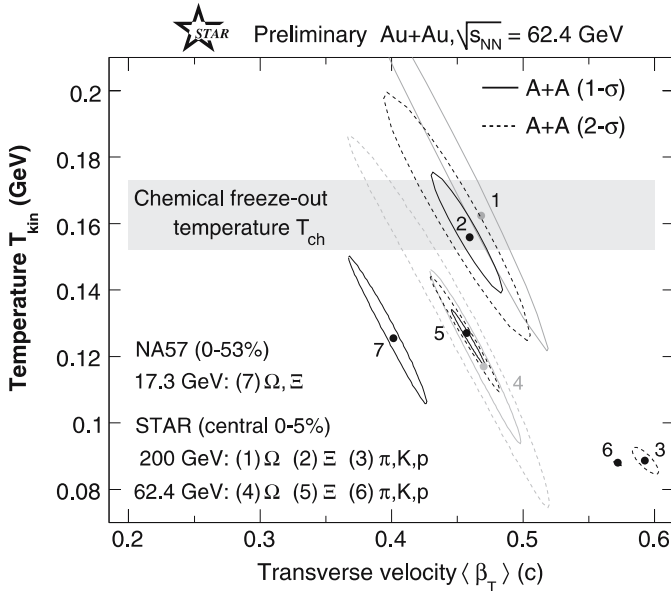


Fig. 4. One and two σ contours in $T_{\text{kin}} - \langle\beta_T\rangle$ space obtained by *Blast-Wave* fits on the measured spectra from most central collisions at different energies and for different particle species. The NA57 contour is from [26]

agreement may also be seen on T_{kin} . The values of T_{kin} seem to be higher for the multi-strange baryons than those obtained by a combined fit for the light particles. Given the large errors, this seems to be especially the case at 200 GeV and is not as obvious at 62.4 GeV.

For a more accurate comparison with ideal hydrodynamics, we use a model by Kolb et al. that numerically solves the hydrodynamical equations. Computations have been published for the energy of 200 GeV [27–29] and as it is publicly available on the web [30], we used it to get predictions for the energy of 62.4 GeV. The applied equation of state is the same as that at 200 GeV and includes a first order phase transition from a QGP to a hadron gas at $T_c = 165$ MeV. The model assumes partial chemical equilibrium in order to reproduce the abundances of the different particle species. Chemical equilibrium is maintained in the QGP phase and a chemical freeze-out occurs at $T_{\text{ch}} = 164$ MeV. The initial entropy and baryon densities have been adjusted to reproduce the measured multiplicity and p/π and \bar{p}/p ratios at 62.4 GeV. This model uses an additional parameter, noted α , that parameterizes a possible initial transverse boost that may have been developed prior to thermalization. The time from which the thermalization hypothesis is valid is parameterized by τ_0 . Its value at 62.4 GeV has been taken to be the same as at 200 GeV [27–29], namely $\tau_0 = 0.6$ fm/c.

Figure 5 shows the spectra for light particles (π^- , K^- , K_S^0 and \bar{p}) for most central Au + Au collisions at 62.4 GeV superposed with hydrodynamical calculations using the aforementioned adapted model. The agreement between data and model holds for a limited domain in transverse

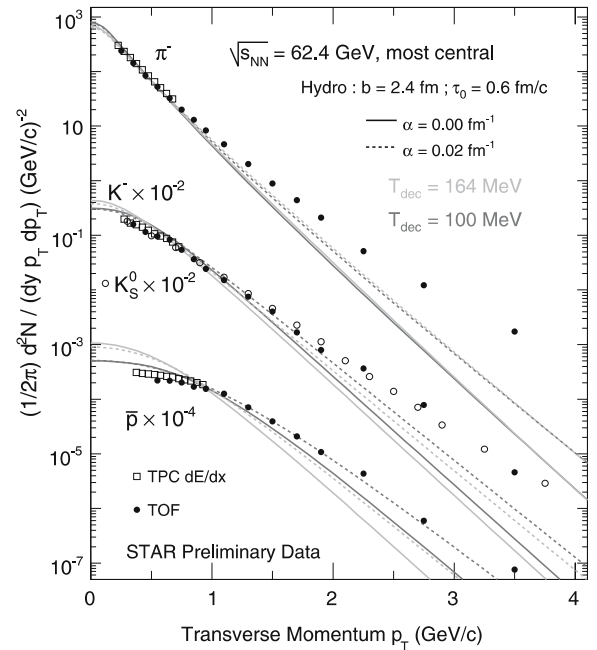


Fig. 5. Measured π^- , K^- , K_S^0 and \bar{p} spectra (*points*) for most central Au + Au collisions at 62.4 GeV superposed with hydrodynamical calculations (*lines*) for different freeze-out temperatures (T_{kin}) and α parameters obtained by using [27–30]

momentum that might be smaller as was observed at 200 GeV [27–29]. The best agreement at 62.4 GeV, as has been at 200 GeV, is obtained with a freeze-out temperature of $T_{\text{kin}} = 100$ MeV and a small but finite value for the α parameter. The agreement between data and hydrodynamical calculations is supposed to hold only at low- p_T (< 1.5 – 2.0 GeV/c), as particles with higher p_T undergo less re-scattering and may therefore verify less the thermalization hypothesis of ideal hydrodynamical calculations. With the current preliminary data, it also looks as if the agreement would hold for higher p_T for the heavier particles. This appears to be true also for the Ω , the particle with the highest strangeness content, where the agreement seems to hold for the entire measured spectra as Fig. 6 illustrates. The non-chemical equilibrium part of the evolution does not include a strange chemical potential. This may have significant influence on the strange particle multiplicity. Therefore we only discuss the shape of the spectra and not its normalization, that has been slightly adjusted to reproduce the data. A similar observation regarding the agreement between data and hydrodynamical calculations on the Ω has already been made at 200 GeV [27–29]. At both energies, it seems that a common freeze-out temperature of ~ 100 MeV and a non-vanishing α parameter, the same as for the light particles, best reproduce the Ω data within the framework of these hydrodynamical model calculations. A higher thermal freeze-out temperature, just after hadronisation ($T_{\text{dec}} = 164$ MeV), does not develop sufficient radial flow to reproduce the data.

The results regarding the thermal freeze-out temperature from these hydrodynamical calculations seem incompatible with the observations from the *Blast-Wave* fits. This apparent discrepancy may have different explanations. First of all, the error bars on the *Blast-Wave* contours (Fig. 4) are still quite large for the multi-strange baryons. The shown contours do not include systematic error induced by the variation of an additional parameter in the *Blast-Wave*, the velocity profile, which affects particularly the values of the temperature parameter. Additionally the fit-range used for the fit on the different particles is not the same, as the measured spectra do not have the same p_T coverage, especially at very low- p_T (< 0.7 GeV/c). Further investigations included a direct comparison between the hydrodynamical (T_{dec}) and the *Blast-Wave* (T_{kin}) temperature. This was done by performing a *Blast-Wave* fit on the spectra from the hydrodynamical model with a given temperature. The fit gives a good agreement with the spectra, however the values of the parameters obtained by the *Blast-Wave* fit are systematically lower than the hydrodynamical temperature with a smaller difference for lower T_{dec} (3–10 MeV at $T_{\text{dec}} \sim 100$ MeV) than at higher T_{dec} (15–30 MeV at $T_{\text{dec}} \sim 164$ MeV). This raises the question whether T_{dec} (hydro) and T_{kin} (B–W) are the same physical quantity.

The 62.4 GeV high statistics data also allows for the study of the centrality dependence of the Ω spectra shown in Fig. 7. The hydrodynamical calculations for the different centrality intervals only differ by the impact parameter, b , used in the calculations. All the other parameters are identical to those used for the most central collisions. Only

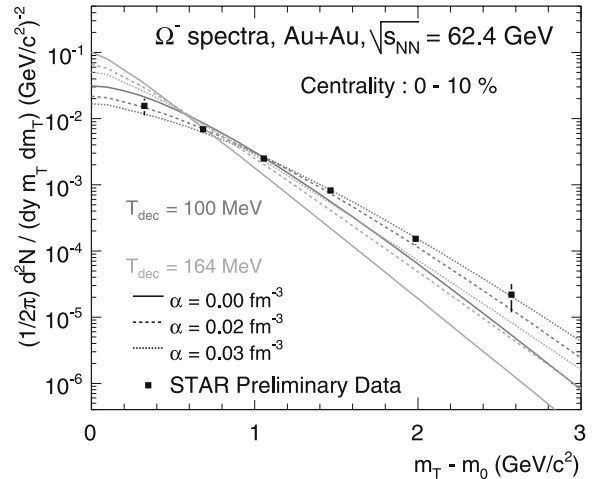


Fig. 6. Measured transverse mass ($m_T = \sqrt{p_T^2 + m_0^2}$) Ω spectra (points) from most central Au + Au collisions at 62.4 GeV superposed with various hydrodynamical calculations (lines) for different T_{dec} and α obtained by using [27–30]

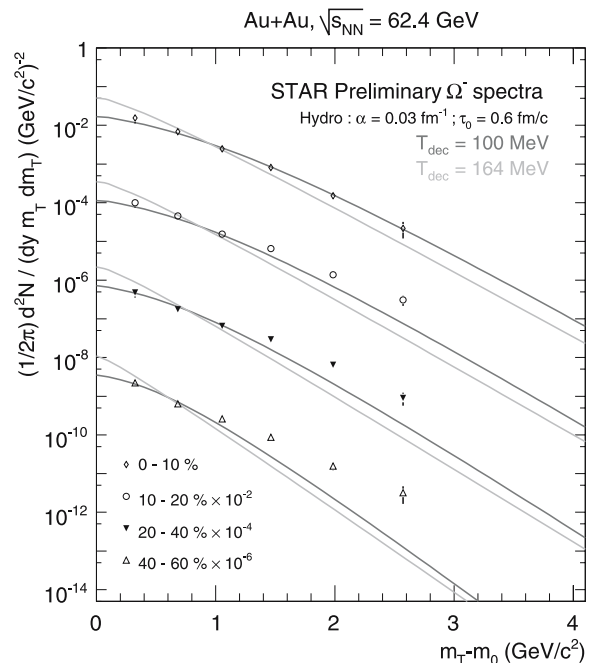


Fig. 7. Centrality dependence of Ω spectra in Au + Au collisions at 62.4 GeV. Superposed with data (points) are hydrodynamical calculations (lines) obtained by using [27–30]

the curves that give the best agreement for the most central collisions (Fig. 6; $\alpha = 0.03 \text{ fm}^{-1}$) are shown for two different temperatures. The spectra from hydrodynamical calculations deviate more and more from the data when the collisions become more peripheral. This deviation is expected as for more peripheral collisions, the system size gets smaller so that the thermalization hypothesis is less valid. The centrality dependence may then help quantify the conditions for which the hypothesis of thermal equilibrium is applicable.

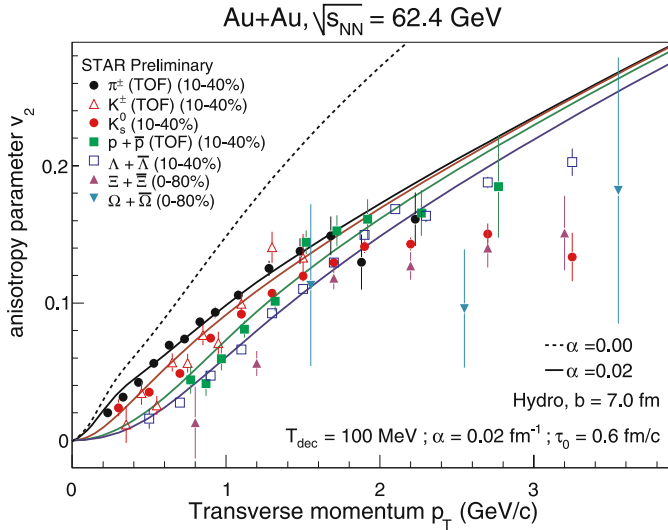


Fig. 8. Elliptic flow with respect to transverse momentum for identified particles in minbias (0%–80%) Au + Au collisions. The data (points) are superposed with hydrodynamical calculations using [27–30]. The dotted line corresponds to the default ($\alpha = 0.00 \text{ fm}^{-1}$; for π^- only) and the full lines are obtained with $\alpha = 0.02 \text{ fm}^{-1}$

Finally, differential elliptic flow behaviour is investigated. Elliptic flow originates from transcription of the spatial asymmetry, caused by the almond shape of the collision overlap zone, into momentum space. Comparisons with data at 200 GeV show that hydrodynamical calculations allow for qualitative reproducibility [31]. The model applied on the 200 GeV data used a complete chemical equilibrium until thermal freeze-out [32]. With this model, it was therefore not possible to reproduce both spectra and elliptic flow. The Kolb et al. model can approximately reproduce the spectra but the default value for the elliptic flow is too large [33].

In Fig. 8, the elliptic flow as a function of transverse momentum is represented for different identified particles in Au + Au collisions at 62.4 GeV. As at 200 GeV, the default value for α (0.00 fm^{-1}) generates too much flow, so that the hydrodynamical calculations overshoot the data (dotted line). The use of $\alpha = 0.02 \text{ fm}^{-1}$ (full lines) allows to simultaneously reproduce spectra and elliptic flow with the same parameters. This reproduction of elliptic flow consists, as at 200 GeV, of the mass ordering at low- p_T and the qualitative agreement of data with hydrodynamical calculations until $p_T \sim 2 \text{ GeV}/c$.

4 Summary and conclusion

The high quality and quantity of data at different energies allow for systematic comparison of physics observables and excitation functions of strange and multi-strange particles. Thermal model fits lead to results at the intermediate energy of 62.4 GeV that are almost identical to the results obtained at 200 GeV, in terms of chemical freeze-out tem-

perature and strangeness saturation factor. The strange and multi-strange anti-baryon to baryon p_T integrated ratios expose a smooth evolution from SPS to top RHIC energies.

Dynamical properties are investigated using a *Blast-Wave* parameterization and an ideal hydrodynamical model description. The agreement of the hydrodynamical models with measured central spectra for light particles may be slightly worse at 62.4 GeV than what was observed with the same model at 200 GeV. Within the framework of this model, the same thermal freeze-out temperature allows a reasonable description of all spectra at the lower and top RHIC energies, including multi-strange baryons. The apparent discrepancy between hydrodynamical model calculations and *Blast-Wave* results may be investigated further with more precise measurements at lower p_T for the multi-strange baryons ($< 0.7 \text{ GeV}/c$). The additional parameter α in the considered hydrodynamical model allows for a simultaneous description of spectra and elliptic flow with the use of an equation of state that involves a phase transition from a QGP to a hadron gas.

Although complete ideal hydrodynamical calculations do not reproduce all the data ($p_T > 2 \text{ GeV}/c$, peripheral, high η , etc.) they give an indication as to what extent the system behaves thermally. In that matter they give the best ever observed agreement with data at RHIC energies. The α parameter in the Kolb et al. model that we used tries a basic description of the pre-equilibrium phase. This parameter as well as the use of a partial chemical equilibrium may to some extent already be considered as extensions to an ideal hydrodynamical description. The whole description may then be obtained with more sophisticated extensions as the combination of ideal hydrodynamics with microscopic cascade calculations [34, 35] or by introducing viscosity [36].

References

1. J. Rafelski, B. Müller, Phys. Rev. Lett. **48**, 1066 (1982)
2. K. H Ackermann et al., Nucl. Instrum. Methods A **499**, 624 (2003)
3. M. Anderson et al., Nucl. Instrum. Methods A **499**, 659 (2003)
4. STAR Collaboration, J. Adams et al., nucl-ex/0606014, submitted to Phys. Rev. Lett.
5. STAR Collaboration, J. Takahashi, J. Phys. G: Nucl. Part. Phys. **41**, S1061 (2004)
6. STAR Collaboration, J. Speltz, J. Phys. G: Nucl. Part. Phys. **41**, S1025 (2004)
7. P. Braun-Munzinger et al., Phys. Lett. B **518**, 41 (2001)
8. F. Becattini et al., Phys. Rev. C **69**, 024905 (2004)
9. J. Cleymans et al., Phys. Rev. C **65**, 027901 (2004)
10. J. Rafelski et al., Phys. Rev. C **72**, 024905 (2005)
11. M. Kaneta, N. Xu, nucl-ex/0405068
12. STAR Collaboration, O. Barannikova, nucl-ex/0403014
13. STAR Collaboration, J. Adams et al., nucl-ex/0311017
14. F. Karsch et al., Eur. Phys. J. C **29**, 549 (2003)
15. J. Cleymans et al., Phys. Rev. C **73**, 034905 (2006)
16. H. Satz, Nucl. Phys. A **715**, 3 (2003)

17. V. Koch, Nucl. Phys. A **715**, 108 (2003)
18. H. Caines et al., J. Phys. G **27**, 311 (2001)
19. F. Antinori et al., Phys. Lett. B **595**, 68 (2004)
20. C. Adler et al., Phys. Rev. Lett. **89**, 092301 (2002)
21. T. Anticic et al., Phys. Rev. Lett. **93**, 022302 (2004)
22. J. Adams et al., Phys. Rev. Lett. **92**, 027901 (2004)
23. J.-Y. Ollitrault, Phys. Rev. D **46**, 229 (1992)
24. F. Retiere, M. Lisa, Phys. Rev. C **70**, 044907 (2004)
25. E. Schnedermann et al., Phys. Rev. C **48**, 2462 (1993)
26. C. Alt et al., Phys. Rev. Lett. **94**, 192301 (2005)
27. P.F. Kolb et al., Phys. Rev. C **62**, 054909 (2000)
28. P.F. Kolb, R. Rapp, Phys. Rev. C **67**, 044903 (2003)
29. P.F. Kolb, U. Heinz, nucl-th/0305084
30. P.F. Kolb, <http://nt3.phys.columbia.edu/OSCAR/models>
31. STAR Collaboration, J. Adams et al., Phys. Rev. C **72**, 014904 (2005)
32. P. Huovinen et al., Phys. Lett. B **503**, 58 (2001)
33. PHENIX Collaboration, K. Adcox et al., Nucl. Phys. A **757**, 184 (2005)
34. T. Hirano, M. Gyulassy, Nucl. Phys. A **769**, 71 (2006)
35. C. Nonaka, S.A. Bass, nucl-th/0607018
36. D. Teaney, Phys. Rev. C **68**, 034913 (2003)